

1           **Two metamorphic cycles recorded by monazite in eclogite-facies**  
2           **gneisses (Southern Armorican Massif, France): A Cambro-Ordovician**  
3           **continental crust involved in eo-Variscan subduction.**

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11          **Abstract:**

12          We present U-Th/Pb data obtained for xenotime and monazite from the polycyclic eclogite-  
13          facies para- and ortho-gneisses of the Les Essarts Unit (Vendée, southern Armorican Massif,  
14          France), which have recorded an HT-LP cycle ending with a first retrogression and a subsequent  
15          HP eclogite-facies metamorphism similar to that of the neighbouring eclogites. Some  
16          paragneisses and orthogneisses are only slightly deformed and retrogressed, showing the  
17          structure of nebulitic migmatite or metagranite, respectively; both show well-preserved  
18          complex coronitic and pseudomorphic microstructures, due to the eclogite-facies  
19          metamorphism. Monazite I and xenotime crystallised during the HT stage providing an  
20          opportunity to date the early HT metamorphic event in the paragneiss and the emplacement of  
21          granite in the orthogneiss. U/Pb ages obtained from monazite I and xenotime of the cordierite-  
22          bearing migmatitic paragneisses range between 510 and 480 Ma (Late Cambrian-Early

23 Ordovician). These ages may correspond to the crystallisation and/or re-equilibration of  
24 monazite and xenotime during the prograde stage of the HT cycle, close to the  $T$  peak.  
25 Consistent monazite and xenotime U/Pb ages around 496 Ma in the orthogneiss represent the  
26 age of the granite protolith. During subsequent HP overprint in the gneisses, numerous coronas  
27 developed at the expense of the early HT parageneses, in particular plagioclase. In both  
28 paragneiss and orthogneiss, monazite I in contact with HT plagioclase reacted to form apatite  
29 + zoisite + monazite II coronas. The small monazite II crystals could be dated in a paragneiss  
30 sample and gave a lower intercept age of  $395 \pm 9$  Ma, interpreted as the age of the eclogite-  
31 facies HP metamorphism. This age is in agreement with those obtained in HP metamorphic  
32 rocks of the Upper Allochthon Unit of the Iberian-Armorican Arc (Bragança, Cabo-Ortegal,  
33 Audierne) representing the first evidence of convergence in the Variscan cycle.

#### 34 **Résumé:**

35 Nous présentons des âges U-Th/Pb obtenus pour le xénotime et la monazite des paragneiss  
36 et orthogneiss polycycliques de l'unité des Essarts (Vendée, domaine sud armoricain, France),  
37 qui ont enregistré un cycle de HT-BP se terminant par une première phase de rétro-morphose,  
38 suivi d'un métamorphisme postérieur dans le faciès éclogite, semblable à celui des éclogites à  
39 proximité. Certains paragneiss et orthogneiss sont très peu déformés et rétro-morphosés,  
40 montrant respectivement la structure initiale d'une migmatite nébuleuse ou d'un métagranite.  
41 Ils présentent tous des textures coronitiques et des pseudomorphoses complexes formées au  
42 cours du métamorphisme en faciès éclogite. La monazite I et le xénotime ont cristallisé pendant  
43 le stade de HT, ce qui permet de dater ce métamorphisme précoce dans le paragneiss, et la mise  
44 en place du granite dans l'orthogneiss. Les âges U/Pb obtenus dans la monazite I et le xénotime  
45 des paragneiss migmatitiques à cordiérite se situent entre 510 et 480 Ma (fin du Cambrien-  
46 début de l'Ordovicien). Cet intervalle d'âges peut correspondre à la cristallisation et/ou au  
47 rééquilibrage de la monazite et du xénotime au cours de la phase prograde du cycle de HT,

48 proche du pic en température. Les âges U/Pb autour de 496 Ma dans la monazite et le xénotime  
49 de l'orthogneiss représentent l'âge du protolithe granitique. Au cours du métamorphisme de HP  
50 ultérieur, de nombreuses couronnes se sont développées aux dépens des paragenèses de HT  
51 précoces, en particulier le plagioclase. Dans les paragneiss et les orthogneiss, la monazite I en  
52 contact avec le plagioclase de HT a réagi pour former des couronnes d'apatite, zoisite et  
53 monazite II. Les cristaux de monazite II de très petite taille ont pu être datés dans un échantillon  
54 de paragneiss et ont donné un âge de  $395 \pm 9$  Ma, interprété comme l'âge du métamorphisme  
55 de HP dans le faciès éclogite. Cet âge est en accord avec ceux obtenus dans les roches  
56 métamorphiques de HP de l'unité Allochtone Supérieure de l'Arc ibéro-armoricain (Bragance,  
57 Cabo-Ortegal, Audierne) qui représentent les premières manifestations de la convergence du  
58 cycle varisque.

59

## 60 **1- Introduction**

61 The record of polycyclic metamorphic evolution is well established in recent collisional  
62 chains such as the Alps, where evidence for earlier Variscan events has long been recognised  
63 (e.g., Le Bayon *et al.*, 2006; Cenki-Tok *et al.*, 2011; Nosenzo *et al.*, 2021). In other contexts,  
64 the preservation of two (or more) orogenic cycles within the rock record is open to debate.  
65 Indeed, polycyclic evolution is not always easy to distinguish from polymetamorphic evolution,  
66 which is characterised by a continuous evolution of *P-T* conditions during a single orogenic  
67 cycle. While structural and petrological studies can help in establishing successive orogenic  
68 events, key evidence often comes from geochronological data, when these events are separated  
69 by a span of time well beyond the estimated duration of an orogenic cycle. Only the ability to  
70 date mineralogical reactions in situ and in their textural context can shed light on this type of  
71 problem, which has large-scale consequences for the evolution of tectono-metamorphic units

72 during orogenesis. The use of geochronometers as faithful petrological and kinematic markers  
73 (hygrochronometers as defined by Villa, 2016) is therefore essential.

74 The Variscan belt of western and central Europe contains eclogite-facies rocks that are  
75 thought to have formed during eo-Variscan subduction. However, estimates of the age of  
76 associated ophiolites and their high-*P* (HP) metamorphism have given widely scattered results  
77 and remain debated. To date, the geodynamic history preceding the subduction events also  
78 remains poorly understood or investigated. Initially, studies aimed at determining the age of HP  
79 events were based on U/Pb dating of zircon yielding widely scattered ages ranging from the  
80 Cambrian to the Lower Silurian (*e.g.*, Gebauer *et al.*, 1981; Peucat *et al.*, 1982; Berger *et al.*,  
81 2010; Whitney *et al.*, 2015). However, it is now clear that zircon evolution is difficult to connect  
82 with the metamorphic history, and clear evidence for metamorphic zircon overgrowths  
83 contemporaneous with the eclogite-facies event is generally lacking (Paquette *et al.*, 2017; Pitra  
84 *et al.*, 2022). Recent studies have re-examined the age of eclogites in the western Variscan belt,  
85 particularly in the French Massif Central (Lotout *et al.*, 2018, 2020; Hoÿm de Marien *et al.*,  
86 2023), using geochronometers such as U-Th/Pb in monazite, U/Pb in rutile or titanite, and  
87 Lu/Hf and/or Sm/Nd in garnet. By combining trace element analysis with in-situ dating of  
88 accessory and major phases, the dates have been interpreted in relation to the petrological  
89 evolution of the sample. The results constrain the age range of the subduction events, and better  
90 define the age of the metamorphic events in this area at around 380–360 Ma. In the Armorican  
91 Massif, ages of HP metamorphism obtained by zircon U/Pb dating also range from 440 Ma to  
92 355 Ma (Peucat *et al.*, 1982; Paquette *et al.*, 1985; Paquette, 1987; see Ballèvre *et al.*, 2009 for  
93 a detailed review). Paquette *et al.* (2017) have re-examined the zircon previously analysed by  
94 isotope dilution thermal ionisation mass spectrometry (ID-TIMS), mostly multigrain fractions,  
95 using the in-situ laser ablation inductively coupled plasma mass spectrometry (LA-ICPMS)  
96 method. Their results led to the conclusion that “no high-pressure event at the previously

97 assumed date of 400–410 Ma is confirmed in any studied zircon grain” and that the “old concept  
98 of a Siluro-Devonian eclogite-facies event is most probably related to analytical limitations of  
99 the published ages in the 1980s, which led to geodynamic overinterpretations” (Paquette *et al.*,  
100 2017). Although the latter study is a fundamental contribution to the interpretation of the zircon  
101 ages obtained by ID-TIMS in HP metamorphic rocks, the ages of the HP metamorphism in the  
102 Armorican Massif obtained using various other methods, such as Sm/Nd Grt ages of  $362 \pm 25$   
103 Ma in the Champtoceaux complex (Bosse *et al.*, 2000) and Rb/Sr and  $^{40}\text{Ar}/^{39}\text{Ar}$  ages of 358–  
104 366 Ma in well-preserved blueschists from the Ile de Groix (Bosse *et al.*, 2005), are still  
105 debated.

106 The discovery of polycyclic eclogite-facies gneisses in Les Essarts Unit (Vendée, southern  
107 Armorican Massif, France: Fig. 1; Godard, 2009) provides an opportunity for a new approach  
108 to HP-metamorphism dating. Petrological evidence has shown that the paragneisses and  
109 orthogneisses of this area have undergone a high- $T$  (HT) low- $P$  (LP) metamorphism, typical of  
110 the continental crust, followed by a first retrogression and then the same HP eclogite-facies  
111 metamorphism as the neighbouring eclogites (Godard, 2009). These gneisses could therefore  
112 represent remnants of a pre-Variscan continental crust (HT migmatites and granites) that were  
113 involved in the eo-Variscan convergence and subduction together with the oceanic crust that  
114 generated the eclogites (HP overprint). In the gneisses, monazite crystallised during the HT  
115 stage and partly recrystallized during the subsequent HP overprint (Godard, 2009), providing  
116 an opportunity to date both events. This article is devoted to the geochronology of these  
117 polycyclic eclogite-facies gneisses, with the aim of dating (a) the metamorphic history of the  
118 continental slice to which they belonged and (b) the eo-Variscan subduction in which they were  
119 involved, prior to their accretion into the Variscan belt.

## 120 **2- Geological setting and petrology**

121 The Les Essarts HP unit occurs in the Southern Armorican Massif (Vendée, Western France),  
122 where it defines a NW–SE-trending zone about 70 km long and a few km wide between two  
123 late-Variscan dextral faults, the Vendée coal belt and the Sainte-Pazanne-Mervent tectonic  
124 lineament (Fig. 1; *e.g.*, Godard, 2001, 2009; Godard et Bonnet, 2007). It comprises altered  
125 ultramafites, eclogites, and eclogite-derived amphibolites, which form several kilometre-long  
126 lenses (Fig. 1) that are stretched and boudinaged parallel to the main vertical foliation in the  
127 surrounding ortho- and para-gneisses.

### 128 **2.1- Eclogites**

129 The Vendée eclogites have been considered to be remnants of an old oceanic crust,  
130 eclogitised during an eo-Variscan subduction and subsequently incorporated into the Variscan  
131 orogenic belt during continental collision (Godard, 1988, 2001, and references therein).  
132 Geochemical studies have shown a rather large mineralogical diversity of eclogites, with (i)  
133 rare zoisite and kyanite-bearing eclogites derived from Al-Mg rich cumulate gabbros, (ii)  
134 abundant quartz-bearing eclogites derived from N-MORB type gabbros, (iii) rutile-rich  
135 eclogites derived from Fe-Ti rich gabbros, and (iv) a few more felsic rocks (Godard, 1988,  
136 2001). Weathered ultramafic rocks with rare remnants of garnet peridotite have also been  
137 observed (Fig. 1; Godard, 2001). Eclogite-facies conditions were first estimated to be  $T = 650$ –  
138  $750^{\circ}\text{C}$  and  $P > 1.6$  GPa (Godard, 1988). Massone and Li (2022) studied an atoll garnet-bearing  
139 eclogite and proposed the existence of two garnet generations reflecting two metamorphic  
140 stages close to the pressure peak (stage 1: 2.0 GPa and  $550^{\circ}\text{C}$  and stage 2: 1.8–2.0 GPa and  
141  $640^{\circ}\text{C}$ ). A major deformation event occurred during the eclogite-facies metamorphism (Godard  
142 and van Roermund, 1995; Mauler *et al.*, 2001), which ended with an almost isothermal  
143 evolution at the beginning of the exhumation (Godard, 2001; Massonne and Li, 2022), followed

144 by various retrograde metamorphic reactions that transformed most of the eclogites into  
145 amphibolites.

146 Geochronological studies of zircon from an eclogite sampled at La Gerbaudière quarry  
147 (Peucat *et al.*, 1982; Peucat, 1983; Postaire, 1983) gave a discordia age of  $436 \pm 15$  Ma (U/Pb  
148 multigrain size-fraction ID-TIMS; lower intercept age), which was then interpreted as the age  
149 of the HP metamorphism. Zircon crystals from the very separate used by Peucat *et al.* (1982)  
150 were analysed in situ by LA-ICPMS dating (Paquette *et al.*, 2017). The results showed the  
151 absence of inherited cores, Th/U ratios of  $0.4 \pm 0.1$  typical of magmatic zircons and yielded a  
152 lower intercept age of  $487 \pm 12$  Ma in a Tera-Wasserburg diagram. This age has therefore been  
153 assigned to the formation of the gabbroic protolith in an oceanic environment, without a record  
154 of the eclogite-facies metamorphism (Paquette *et al.*, 2017). Some of the eclogites were  
155 exhumed at the end of the Carboniferous (at about 300 Ma), since eclogite pebbles have been  
156 observed in the Stephanian detrital deposits of the Vendée coal belt (Godard, 2001; Godard et  
157 Bonnet, 2007).

## 158 **2.2- Coronitic gneisses**

159 Most of the rocks surrounding the eclogites are schistose and rich in white mica. Their  
160 present paragenesis consists of phengitic muscovite, quartz, albitic plagioclase, biotite and,  
161 sporadically, garnet and microcline. However, relict parageneses suggest that they are derived  
162 from earlier biotite, garnet, quartz, plagioclase and microcline-bearing gneisses, with cordierite  
163 in some cases. In a dozen localities listed by Godard (2001), some of these gneisses are  
164 exceptionally poorly deformed and retrogressed (Fig. 2). Early parageneses and coronitic  
165 structures are then preserved, revealing a complex history.

166 Two types of gneisses can be distinguished on the basis of their origin:

167 (a)- *Orthogneisses* show the usual paragenesis of a metagranite (Fig. 2b): quartz +  
168 oligoclase + biotite + muscovite + microcline. Orthoclase is transformed into perthitic  
169 microcline, which can form centimetre-scale *augen*. The main feature of these orthogneisses is  
170 the presence of complex garnet-bearing microcrystalline coronas along the plagioclase-biotite  
171 and microcline-biotite interfaces.

172 (b)- *Paragneisses* also show numerous coronas that have developed at the expense of a  
173 pre-existing cordierite-bearing migmatitic paragenesis (Fig. 2a). The *P-T* conditions of this HT-  
174 LP event have been estimated to be ~0.32 GPa and ~670°C in samples from Grezay (Godard,  
175 2009). Commonly, the structure of the paragneisses is that of a nebulitic migmatite (Fig. 2a). In  
176 the Grezay area, migmatitisation is also evidenced by the common occurrence of leucosomes  
177 composed of Qz + Pl + Kfs + Bt ± Crd (± tourmaline ± Ilm ± monazite). The melanosome  
178 generally includes Bt + Qz + Crd + Grt + Pl + Ilm (+ F-apatite + monazite), with relics of  
179 muscovite and Al-silicate isolated within biotite-rich clusters and pseudomorphs after  
180 cordierite, respectively. Biotite, quartz and Al-silicate are always separated from each other by  
181 pseudomorphed cordierite which indicates that cordierite grew at the expense of biotite, quartz  
182 and sillimanite, according to a well-known migmatitisation reaction:  $Bt + Qz + Sil (\pm Pl) \rightarrow Crd$   
183  $+ Melt (+ Grt)$  (*e.g.*, Waters and Whales, 1984; Le Breton and Thompson, 1988; Stevens *et al.*,  
184 1997).

185 Numerous coronas can be observed in both types of rock (Fig. 2d). They developed at the  
186 expense of the pre-existing migmatitic or granitic parageneses, particularly at the interfaces  
187 between feldspars (plagioclase or microcline) and mafic minerals (biotite, ilmenite or  
188 cordierite: Fig. 2c-d), but also monazite. Plagioclase commonly forms a microcrystalline  
189 mosaic of polycrystalline albite in which minute rodlets of kyanite, and minor zoisite and micas  
190 are visible only by scanning electron microscopy (SEM). The kyanite rods are ordinarily  
191 arranged in trails that delineate polygonal millimetre-sized cells with a honeycomb-like



192 structure (Fig. 2c), which is interpreted by Godard (2009) as the silhouettes of the former  
193 plagioclase single crystals of the HT paragenesis. This type of recrystallisation without  
194 deformation would be explained by phase transitions: Oligoclase  $\rightarrow$  X + minor zoisite and  
195 kyanite, followed by the retrograde reaction X  $\rightarrow$  albite; indeed, X could be jadeite, which,  
196 however, has never been observed. This second episode occurred in eclogite-facies conditions  
197 at  $P > 1.6$  GPa and  $T \approx 700^\circ\text{C}$  (Godard, 2009).

198 Two lines of evidence indicate that the HT event and the subsequent HP eclogite-facies  
199 overprint were separated by a first retrogression at  $T < 400^\circ\text{C}$ : K-feldspar exsolved perthitic  
200 lamellae and cordierite was altered to pinite before the formation of the HP coronas, indicating  
201 a polycyclic rather than polymetamorphic  $P$ - $T$  evolution for these rocks.

202 Although the eclogites and ultramafites have been considered to be remnants of an oceanic  
203 lithosphere, the surrounding ortho- and para-gneisses clearly belong to a continental crust. This  
204 led Godard (1983, 2001) to interpret the Les Essarts Unit as a possible tectonic melange of pre-  
205 Hercynian continental and oceanic crusts that were eclogitised during eo-Variscan subduction  
206 and subsequently incorporated into the Hercynian orogenic belt.

207 Two samples of the paragneisses from Grezay were studied in the present work: VP7-2 is a  
208 piece of the main sample described by Godard (2009), whereas VP7c was sampled in the same  
209 disused small quarry. VP7r is an orthogneiss (actually, an almost undeformed metagranite: Fig.  
210 2b) from the same locality (lat. 46.7820; long. -1.2597; Fig. 1).

### 211 **3 – Monazite and xenotime**

#### 212 *3.1 Textural characteristics and chemical composition of monazite*

213 In the three samples studied, monazite is common and occurs either as inclusions in garnet,  
214 quartz or biotite or at contacts with the plagioclase pseudomorphs described above. Monazite

215 isolated in garnet, quartz or biotite (partially transformed into phengite) consists in pristine  
216 prismatic or rounded grains of 10–30  $\mu\text{m}$  in size (Figs. 3e-g, 4a-c, 4e-f). They appear not to  
217 have chemically interacted with the host mineral, but show  $\sim 20$   $\mu\text{m}$ -thick cathodoluminescent  
218 halos in quartz, visible under the SEM, revealing lattice damage due to monazite radioactivity.  
219 Monazite I crystals at contacts with the plagioclase pseudomorphs commonly reach 50  $\mu\text{m}$  in  
220 size, and show complex coronas composed of apatite, zoisite and clusters of small grains of  
221 monazite II, both in orthogneiss (sample VP7r: Fig. 5) and in paragneiss (sample VP7-2: Fig.  
222 3a-d). In sample VP7r (Fig. 5), the monazite II crystals are small ( $< 5$   $\mu\text{m}$ ) compared to sample  
223 VP7-2 where their size can reach 20  $\mu\text{m}$  (Fig. 3a-d). In the paragneiss sample VP7c, where  
224 large ( $> 100$   $\mu\text{m}$ ) monazite I grains are commonly subhedral, elongated or round, monazite II  
225 crystals show irregular corroded rims at the contacts with the plagioclase pseudomorphs (Fig.  
226 4d), but no clear reaction corona. By contrast, monazite I grains are completely replaced by  
227 clusters of monazite II associated with apatite and zoisite in the other two samples, in particular  
228 VP7-2 (Fig. 3c, d).

229 The chemical compositions of monazite are reported in Supplementary Materials (SM1) and  
230 in Figure 6a-b. In all samples, there is no systematic relationship between monazite  
231 compositions and textural position. Most of the observed chemical variations are explained by  
232 the brabantite substitution ( $2 \text{REE}^{3+} = \text{Th}^{4+} + \text{Ca}^{2+}$ ) (Fig. 6a), with some huttonite exchange  
233 ( $\text{REE}^{3+} + \text{P}^{5+} = \text{Th}^{4+} + \text{Si}^{4+}$ ). High-contrast BSE imaging and X-ray mapping (SM5) show that  
234 most of the monazite grains, either enclosed in quartz, garnet, micas or in the pseudomorphs  
235 after plagioclase, are only slightly zoned, with Y, U and Th showing the most variation (Fig.  
236 6b). However, it is difficult to define a general trend for the core-to-rim zonation of the grains.

237 Pristine monazites in the studied paragneisses (*i.e.*, monazite in sample VP7c and non-  
238 reactive monazite in sample VP7-2) show large variations in  $\text{Y}_2\text{O}_3$  content (Fig. 6b), with  
239 monazites from sample VP7-2 having a higher  $\text{Y}_2\text{O}_3$  content ( $0.42 < \text{Y}_2\text{O}_3 < 2.02$  wt%) than

240 sample VP7c ( $0.11 < Y_2O_3 < 1.75$  wt%). Monazite therefore appears to be more saturated in Y  
241 in the case of VP7-2, where it coexists with xenotime, as opposed to VP7c where xenotime is  
242 absent. Variations in  $ThO_2$  content are also large, with a slightly wider range and higher content  
243 in sample VP7-2 ( $2.28 < ThO_2 < 5.49$  wt%) than in sample VP7c ( $2.96 < ThO_2 < 4.94$  wt%). In  
244 sample VP7-2, the  $ThO_2$  content decreases as  $Y_2O_3$  increases, whereas in sample VP7c both are  
245 correlated. Monazite included in garnet and quartz in sample VP7-2 shows large variations in  
246  $ThO_2$  content (between 1.49 and 7.79 wt%) and  $UO_2$  content (between 0.31 and 1.87 wt%) (Fig.  
247 6b).  $Y_2O_3$  is slightly higher in the reacted monazite than in the pristine monazite (between 1.04  
248 and 2.58 wt%). In sample VP7c, the monazite inclusions have a chemical composition almost  
249 similar to that of pristine monazites, except for the  $Y_2O_3$  which is slightly higher (2.12 wt%) in  
250 the inclusions. In both samples, the inclusions open to the matrix display the same range of  
251 compositions as the shielded inclusions. In sample VP7-2, the core of the reactive monazites  
252 has  $Y_2O_3$ ,  $ThO_2$  and  $UO_2$  contents between those of the inclusions and those of the pristine  
253 matrix monazite (Fig. 6b). In sample VP7-2, monazite II in the coronas has compositions almost  
254 similar to those of pristine monazite in the same samples. A few monazite I grains of the  
255 paragneiss samples show inherited nuclei, with a distinct composition (see VP7C-45 and VP7-  
256 2-12 in SM5).

257 Monazite from the orthogneiss sample VP7r has a distinctly different composition from that  
258 of the paragneiss samples (Fig. 6a and b). The  $ThO_2$  and  $UO_2$  contents in the core of the reactive  
259 monazite are in the range of 3.6–4.7 wt% and 0.3–1.8 wt% respectively. The maximum  $Y_2O_3$   
260 content is significantly higher than in the other samples (up to 3.8 wt%; Fig. 6b). Only one  
261 analysis of monazite II gave acceptable results (SM1). It is characterised by a low  $ThO_2$  content  
262 and relatively high  $UO_2$  and  $Y_2O_3$  contents (Fig. 6b).

263 In conclusion, variations in monazite composition observed between the different samples  
264 are mainly related to the protolith (paragneiss versus orthogneiss). The different chemical

265 compositions of the monazite crystals in each sample are not clearly related to their textural  
266 positions. It is therefore impossible to distinguish the different generations of monazite in  
267 samples VP7-2 and VP7r on the basis of their chemical composition. Similarly, the internal  
268 zoning observed in pristine monazite or in the core of the reactive monazite is highly variable  
269 from grain to grain (SM5) and is not related to their textural position.

270 The distribution of REE in the monazite of paragneiss sample VP7-2 provides more  
271 discriminating information (Fig. 6c and SM3). Pristine monazites (shielded in quartz) have a  
272 negative Eu anomaly ( $\text{Eu}/\text{Eu}^* \approx 0.13\text{--}0.16$ ), and are 10 times more enriched in HREE compared  
273 to the core of the reactive monazite I (Eu anomaly  $\approx 0.36$  to 0.58). One analysis of monazite II  
274 shows one of the weakest Eu anomalies ( $\text{Eu}/\text{Eu}^* \approx 0.58$ ) and one of the most depleted HREE  
275 content (Fig. 6c).

### 276 *3.2 Textural characteristics of xenotime*

277 Xenotime ( $\text{YPO}_4$ ) is present in two samples (VP7-2 and VP7r). In the orthogneiss sample  
278 VP7r, it occurs as 30–50  $\mu\text{m}$  subhedral grains (Fig. 7e, f), with clear evidence that this accessory  
279 mineral belongs to the granite magmatic paragenesis. In the paragneiss sample VP7-2, xenotime  
280 grains ( $<30 \mu\text{m}$ ) and numerous  $\mu\text{m}$ -sized xenotime inclusions are scattered in large garnet  
281 crystals (Fig. 7b-d), which Godard (2009) interpreted as belonging to the HT paragenesis  
282 (Garnet I). In contrast the Ca-rich garnet coronas and overgrowths (Garnet II, Fig. 7a) that  
283 formed at the contacts with feldspars during the eclogite-facies metamorphism are devoid of  
284 such xenotime inclusions (Fig. 7a-b). Electron microprobe analyses show that the first  
285 generation of garnet (Garnet I), coexisting with xenotime, is relatively rich in Y ( $745 \pm 155 \mu\text{g/g}$   
286 [ppm] Y; 145 spots), which commonly occurs as solid solution of the YAG end-member  
287 ( $\text{Y}_3\text{Al}_2[\text{Al}_3\text{O}_{12}]$ ) in garnet. By comparison, the eclogite-facies xenotime-free Garnet II is much  
288 poorer in it ( $434 \pm 365 \mu\text{g/g}$  Y; 49 spots; Fig.7 b). This suggests an early crystallization of the

289 xenotime in this sample, coeval with Garnet I. The largest xenotime grains could have been  
290 incorporated into Garnet I during its growth, whereas the  $\mu\text{m}$ -sized inclusions could have been  
291 exsolved during its HT prograde evolution, as the solubility of Y in garnet decreases with  
292 temperature (Pyle and Spear, 2000).

#### 293 **4- Monazite and xenotime U-Th/Pb dating**

294 Monazite and xenotime were analysed in thin section by LA-ICPMS in Clermont-Ferrand  
295 (France) for U-Th/Pb dating (Figs. 8 to 10). The analytical methods are presented in the  
296 electronic supplementary material SM3. The results are presented in Tera-Wasserburg diagrams  
297 and U-Th/Pb diagrams (Figs. 8 to 10). All data are reported in SM2 with  $2\sigma$  uncertainties.

298 In monazite, three isotope systems can be used to obtain dates, presented here in order of  
299 decreasing reliability:  $^{208}\text{Pb}/^{232}\text{Th}$ ,  $^{206}\text{Pb}/^{238}\text{U}$  and  $^{207}\text{Pb}/^{235}\text{U}$ .  $^{232}\text{Th}$  is so abundant (here between  
300 1.5 and 7 wt%; SM1) that  $^{208}\text{Pb}$  originated from common Pb can be considered negligible in  
301 most cases compared to radiogenic  $^{208}\text{Pb}$ . It also enables small spots to be analysed by laser  
302 ablation. To ensure that common Pb contamination or isotope disturbance of the Th/Pb system  
303 are minimal (Didier *et al.*, 2013),  $^{208}\text{Pb}/^{232}\text{Th}$  dates can be compared with  $^{206}\text{Pb}/^{238}\text{U}$  dates in  
304  $^{208}\text{Pb}/^{232}\text{Th}$  vs.  $^{206}\text{Pb}/^{238}\text{U}$  diagrams,  $^{238}\text{U}$  being the most abundant isotope of U (typically less  
305 than 1 wt% in our samples). On the other hand, U/Pb systems are sensitive to common Pb  
306 contamination and must be plotted in Tera-Wasserburg diagrams to obtain lower intercept dates  
307 consistent with the common Pb composition. In this study, we always compare the results  
308 obtained in the three systems and consider as the most geologically significant those results that  
309 are consistent in both the  $^{208}\text{Pb}/^{232}\text{Th}$ - $^{206}\text{Pb}/^{238}\text{U}$  and  $^{206}\text{Pb}/^{238}\text{U}$ - $^{207}\text{Pb}/^{235}\text{U}$  systems.

##### 310 *4.1 Paragneiss sample VP7-2*

311 In sample VP7-2, 6 xenotime grains were analysed for a total of 9 analyses retained after  
312 treatment. In this sample, xenotime is present as inclusions in large garnet crystals belonging to  
313 the HT paragenesis or close to the dissolved rims of these crystals (Fig. 7c-d). These inclusions  
314 are either large (30–40  $\mu\text{m}$ ) or  $\mu\text{m}$ -sized, the latter possibly resulting from Y exsolution (see  
315 3.2). Only the first category was large enough to be analysed (Fig. 8), and only one grain was  
316 analysed in the matrix, outside the garnet. In a Tera-Wasserburg diagram, 3 analyses are  
317 discordant and the concordant data range between 516 and 415 Ma. The oldest data are  
318 measured in shielded xenotime inclusions in garnet, defining a weighted mean  $^{206}\text{Pb}/^{238}\text{U}$  date  
319 of  $509 \pm 8$  Ma (SM2; MSWD = 0.51; N = 3).

320 In the same sample, 37 monazite grains were analysed for a total of 106 analyses retained  
321 after treatment (63 in monazite II + 43 in monazite I). For the sake of clarity, we show the data  
322 obtained from monazites I (*i.e.*, core of reactive monazite and non-reactive pristine monazite  
323 included in garnet and quartz; Fig. 9a, b) separately from those obtained from monazite II (*i.e.*,  
324 monazite formed in the corona during reaction; Fig. 9c, d). For monazite I, 6 analyses show  
325 common Pb contamination, whereas 37 concordant analyses in the Tera-Wasserburg diagram  
326 (Fig. 9a) range from 542 to 450 Ma (Fig. 3e-g), the scatter of which prevents the calculation of  
327 a concordia date. Twenty-four concordant  $^{206}\text{Pb}/^{238}\text{U}$ - $^{208}\text{Pb}/^{232}\text{Th}$  dates spread from 541 to 471  
328 Ma. Twenty analyses are concordant in both the U-Th/Pb and the U/Pb systems and provide a  
329 weighted mean  $^{206}\text{Pb}/^{238}\text{U}$  date of  $492 \pm 4$  Ma (Fig. 9a and SM2; mean square weighted  
330 deviation [MSWD] = 2.2; N = 19; 1 data at ~540 Ma excluded) and a weighted mean  $^{208}\text{Pb}/^{232}\text{Th}$   
331 date of  $489 \pm 6$  Ma (MSWD = 3.6; N = 19; see Fig. 9b for the  $^{208}\text{Pb}/^{232}\text{Th}$  date distribution).  
332 Considering only the pristine monazite (*i.e.*, excluding the core of the reacting monazite grains),  
333 the results are similar in the error bar: weighted mean  $^{206}\text{Pb}/^{238}\text{U}$  date of  $490 \pm 6$  Ma (MSWD  
334 = 1.9; N = 11) and weighted mean  $^{208}\text{Pb}/^{232}\text{Th}$  date of  $487 \pm 9$  Ma (MSWD = 4.2; N = 11).

335 Most of the monazite II in the reaction corona around the monazite I gives discordant U/Pb  
336 dates. The discordance is explained by the small grain size (between 2 and 20  $\mu\text{m}$ ) relative to  
337 the laser spot size (7  $\mu\text{m}$ ), which induces a variable amount of mixing with the adjacent phases  
338 (Fig. 3a-d), and the incorporation of various amount of common Pb. Only 13 out of 63 analyses  
339 are concordant in both the U/Pb system and the  $^{206}\text{Pb}/^{238}\text{U} - ^{208}\text{Pb}/^{232}\text{Th}$  system, giving dates  
340 between 558 and 388 Ma (Fig. 9c) and a  $^{208}\text{Pb}/^{232}\text{Th}$  distribution ages defining 2 peaks at about  
341 486 and 398 Ma (Fig. 9d). In a Tera-Wasserburg diagram, 41 analyses define a linear trend  
342 intercepting the concordia curve at  $395 \pm 9$  Ma (Fig. 9c; MSWD = 3.1) with a  $^{207}\text{Pb}/^{206}\text{Pb}_0$  at  
343  $0.93 \pm 0.27$  consistent with the common Pb composition at 395 Ma according to Stacey and  
344 Kramers (1975). Five concordant data between 405 and 388 Ma also allow to calculate a U/Pb  
345 concordia date of  $394 \pm 13$  Ma (insert of Fig. 9c). The same data give a Th-U/Pb concordia date  
346 of  $396 \pm 6$  Ma.

#### 347 *4.2 Paragneiss sample VP7c*

348 Thirty-seven monazite grains were analysed in sample VP7c (Fig. 4) with 37 analyses  
349 retained after treatment. Seventeen analyses show common Pb contamination whereas 20  
350 concordant analyses in the Tera-Wasserburg diagram (Fig. 9e) range from 505 to 462 Ma. This  
351 spread prevents the calculation of a concordia date. Of these data, 18 are also concordant in the  
352  $^{206}\text{Pb}/^{238}\text{U} - ^{208}\text{Pb}/^{232}\text{Th}$  system. In a Tera-Wasserburg diagram, 25 data defined a linear trend  
353 intercepting the concordia curve at  $479 \pm 4$  Ma (MSWD = 1.7;  $^{207}\text{Pb}/^{206}\text{Pb}_0 = 0.46 \pm 0.27$ ; Fig.  
354 9e). The  $^{208}\text{Pb}/^{232}\text{Th}$  date distribution confirm the spread of the data and shows two pics at  
355 around 495 and 480 Ma (Fig. 9f). The oldest dates in this sample between 510 and 504 Ma were  
356 measured in shielded monazite contained in HT garnet crystals (Fig. 4a, SM2).

#### 357 *4.3 Orthogneiss sample VP7r*

358 Ten xenotime grains were analysed in sample VP7r for a total of 12 analyses retained after  
359 treatment (Fig. 10b). In this sample, xenotime was observed in the matrix as rounded grains 40  
360 to 50  $\mu\text{m}$  in size (Fig. 7e, f). Three analyses give concordant U/Pb dates between 506 and 440  
361 Ma, the other data showing various amount of common Pb contamination. Five data among the  
362 oldest define a linear trend that intercepts the concordia curve at  $496 \pm 13$  Ma (MSWD = 2.3;  
363  $^{207}\text{Pb}/^{206}\text{Pb}_0 = 1.25 \pm 0.18$ ; Fig. 10b).

364 Ten monazite grains from the same sample, representing 9 analyses retained after treatment,  
365 were measured (Fig. 10a). All but one of the analyses show a common Pb contamination. The  
366 concordant data (also concordant in the U-Th/Pb systems) gives a  $^{206}\text{Pb}/^{238}\text{U}$  date of  $489 \pm 9$   
367 Ma (Fig. 5c). In a Tera-Wasserburg diagram, 8 data define a linear trend that intercepts the  
368 concordia curve at  $496 \pm 10$  Ma (MSWD = 2.4;  $^{207}\text{Pb}/^{206}\text{Pb}_0 = 0.96 \pm 0.52$ ; Fig. 10a).

369

## 370 **5- Discussion**

### 371 *5.1 Interpretation of the measured dates*

372 In the orthogneiss sample VP7r, the concordant dates obtained for the xenotime grains range  
373 up to 506 Ma. Their intercept date at  $496 \pm 13$  Ma (Fig. 10b), although poorly constrained, is  
374 similar to that obtained for the monazite I grains at  $496 \pm 10$  Ma (Fig. 10a), with which they are  
375 occasionally associated (Fig. 5a). The age of about 496 Ma is therefore almost certainly that of  
376 the granite protolith. This age is close to that obtained for La Roche-aux-Lutins orthogneiss  
377 (Rocheservière, Fig. 1) which also belongs to the Les Essarts unit and yielded an Ordovician  
378 protolith age of  $483 \pm 4$  Ma (ID-TIMS U-Pb zircon, Lahondère *et al.*, 2009).

379 In the VP7-2 paragneiss, xenotime is almost exclusively present in the relatively Y-rich  
380 garnet crystals that have grown during the HT cycle (Garnet I; see section 3.2). The xenotime  
381 dates obtained, despite the small number of analyses, suggest 2 populations at about 509 and



382 480 Ma (Fig. 8). The xenotime-garnet geothermometer of Pyle and Spear (2000) applied to  
383 Garnet I yields temperatures of  $534 \pm 64^\circ\text{C}$ , in agreement with those of the Garnet I growth  
384 estimated by Godard (2009) for the same sample ( $540 \rightarrow 570^\circ\text{C}$  at 0.4 GPa; see A $\rightarrow$ B in Fig. 13  
385 of Godard, 2009). Since the solubility of Y in garnet decreases with increasing  $T$  (Pyle and  
386 Spear, 2000), the evolution towards the  $T$  peak up to about  $670^\circ\text{C}$  (D in Fig. 13 of Godard,  
387 2009) should result in Y exsolution. Exsolution of Y may have produced the numerous  $\mu\text{m}$ -  
388 sized xenotime inclusions in Garnet I (Figs. 7b, d and 4b), which are unfortunately too small to  
389 be properly analysed, and possibly some overgrowth of the pre-existing xenotime inclusions. It  
390 follows that the oldest xenotime age of about 509 Ma provided by the VP7-2 paragneiss should  
391 be assigned to the prograde stage of the HT cycle, whereas the second age at about 480 Ma  
392 could be related to a readjustment close to the  $T$  peak (Fig. 11).

393 The different types of monazite found in the studied samples are related to two main episodes  
394 of monazite growth, namely monazite I of Cambro-Ordovician age, similar to the xenotime  
395 ages obtained in the same samples, and monazite II of Middle Devonian age (see section 4). In  
396 fact, most of the minerals from the initial HT or granitic paragenesis (i.e., biotite, ilmenite,  
397 garnet I, cordierite, monazite I) reacted with the plagioclase pseudomorphs during the eclogite-  
398 facies metamorphism to produce various coronas (Godard, 2009). Monazite I is simply  
399 considered to be one of those minerals that reacted with plagioclase to form, in this case, a  
400 corona of apatite + zoisite + monazite II. An important aspect of these HP coronitic reactions is  
401 the complete transformation of the HT-LP Ca-rich anorthite end-member of plagioclase into a  
402 calcic mineral stable at high pressure. This mineral is a grossular-rich garnet II at the contacts  
403 of plagioclase with biotite, ilmenite or garnet I (Figs. 2d, 7a; Godard 2009). At the monazite-  
404 plagioclase interface, garnet can hardly crystallise due to the lack of Fe and Mg in both  
405 reactants, so apatite and zoisite act as the calcic minerals capable of recycling Ca at high  
406 pressure. Although the  $P$ - $T$  conditions for the formation of this specific corona, unlike other HP

407 coronas, cannot be properly quantified, it is likely to be related to the HP event. Thus, the  
408 formation of monazite I and monazite II would coincide with the two main HT and HP events,  
409 separated by a first retrogression (Godard, 2009), and therefore related to two different orogenic  
410 cycles with a time gap of about 100 Ma.

411 In the three samples, the Th/Pb and U/Pb dates obtained for monazite I are scattered. Some  
412 rare analyses of monazite I (and, incidentally, of monazite II) from paragneiss sample VP7-2  
413 gave dates close to 540–570 Ma (Fig. 9b, d). Assuming that these old ages are geologically  
414 significant, they could represent a fraction of detrital monazite inherited from an earlier orogen,  
415 Cadomian or Pan-African, since the host rock is a metasediment. Such inheritance is supported  
416 by the presence of zones of peculiar composition within certain monazite crystals (*e.g.*, VP7-2-  
417 12 in SM5). The other dates of monazite I differ from sample to sample: ~490 and ~480 Ma for  
418 paragneiss samples VP7-2 and VP7c, and ~496 Ma for the orthogneiss VP7r, although they are  
419 very close within the error bar. In the paragneiss sample VP7-2, textural observations and  
420 chemical compositions of the pristine monazite grains do not support the existence of multiple  
421 populations among the oldest monazite grains. In the paragneiss sample VP7c, the oldest dates  
422 (between 501 and 505 Ma) were obtained for monazite I inclusions armoured in HT garnet I  
423 (Fig. 4a), whereas slightly younger ones were obtained for matrix grains (Fig. 4c-f). This can  
424 hardly be explained by mixed dates with younger monazite II-bearing domains, since this  
425 sample does not show clear reaction coronas producing monazite II (Fig. 4c, d and f). On the  
426 other hand, it may reflect the early trapping of the oldest monazite grains in garnet I, whose  
427 growth took place during the prograde stage of the HT metamorphic cycle, at conditions close  
428 to 550°C (A→B in Fig. 13 of Godard, 2009). Finally, considering the statistical distribution of  
429 pristine monazite Th/Pb dates (concordant in both the  $^{206}\text{Pb}/^{238}\text{U}$ - $^{208}\text{Pb}/^{232}\text{Th}$  and  $^{206}\text{Pb}/^{238}\text{U}$ -  
430  $^{207}\text{Pb}/^{235}\text{U}$  systems), two peaks appear at similar values, around 495 and 480 Ma, in both  
431 paragneiss samples VP7-2 (Fig. 9b) and VP7c (Fig. 9f), in contrast to the unimodal distribution

432 for monazite I from the VP7r orthogneiss metagranite ( $496 \pm 10$  Ma). This suggests that the age  
433 of monazite I, which is well defined for the orthogneiss ( $496 \pm 10$  Ma), shows a greater spread  
434 in the case of the paragneisses (from around 510 to 450 Ma). This spread may indicate monazite  
435 crystallization and/or re-equilibration over a longer period during the prograde stage of the HT  
436 cycle, close to the  $T$  peak. This interpretation is also supported by the spread of xenotime ages  
437 from the VP7-2 paragneiss (Fig. 8).

438 Monazite II could only be analysed in paragneiss sample VP7-2 due to its small size in  
439 sample VP7r. Most of monazite II yields discordant U-Pb analyses because of mixing with Th  
440 and U-poor adjacent phases and incorporation of common Pb during the eclogite-facies  
441 metamorphic event. Considering only the concordant data, the Th-U/Pb dates are scattered  
442 between 507 and 391 Ma (the concordant date of 562 Ma is not considered geologically  
443 significant). Concordant  $^{208}\text{Pb}/^{232}\text{Th}$  distribution dates clearly defined 2 peaks at about 486 and  
444 399 Ma (Fig. 9d). This means that in the metamorphic corona surrounding the monazite I, the  
445 measured dates cover the full range of ages from this sample. In other words, some monazite II  
446 domains appear to have preserved Cambro-Ordovician ages. The REE profile obtained for a  
447 grain of monazite II from sample VP7-2 shows a smaller Eu anomaly and lower HREE contents  
448 compared to monazite I in the same rock, which could reflect a crystallisation of monazite II  
449 coeval with feldspar breakdown and some garnet growth. A lower intercept date at  $395 \pm 9$  Ma  
450 could be obtained with 41 analyses in monazite II and a  $^{207}\text{Pb}/^{206}\text{Pb}$  initial ratio of  $0.93 \pm 0.27$ ,  
451 a value consistent with that of common Pb at 395 Ma, according to Stacey and Kramers (1975).  
452 Moreover, concordant data allow the calculation of a concordia age of  $394 \pm 13$  Ma, consistent  
453 in the 3 isotopic systems (see section 4.1). We interpret the age of  $395 \pm 9$  Ma as the age of the  
454 eclogite-facies event.

## 455 *5.2 Cambro-Ordovician HT-LP cycle*

456 Godard (2009) has described the petrological characteristics of a HT-LP cycle that resulted,  
457 in the Grezay area, in the migmatisation of cordierite-bearing paragneiss and the intrusion of  
458 granite. The evolution of the migmatite indicates a clockwise  $P$ - $T$  path for the HT migmatitic  
459 metamorphism, with  $T$ -peak conditions of about 0.30–0.35 GPa and 650–700 °C (Fig. 11). The  
460 data obtained for monazite I and xenotime, although scattered, allow this HT event to be dated  
461 to around 510–480 Ma (Late Cambrian-Early Ordovician). As discussed above, this scatter may  
462 represent part of the prograde evolution up to the metamorphic peak in the paragneiss, whereas  
463 the same minerals give a more precise age of  $496 \pm 10$  Ma for the orthogneiss, which must be  
464 attributed to crystallisation of the granite protolith.

465 During the Cambro-Ordovician, the Armorican Massif lay along the northern margin of  
466 Gondwana and was subject to continental extension during the early stages of the Gondwana  
467 breakup. Direct evidence of rifting includes widespread emplacement of alkaline/peralkaline  
468 magmas (granitoids and volcanics) at ca. 490–480 Ma in Brittany, Limousin and in Galicia  
469 (Ballèvre *et al.*, 2012; Kroner & Romer, 2013). To explain this Early Ordovician magmatism,  
470 Ballèvre *et al.* (2012) proposed a schematic tectonic evolution that favours the opening of an  
471 intracontinental rift located away from the volcanic arc and its associated back-arc basin (the  
472 Rheic Ocean). In this model, HT metamorphism and partial melting can occur in the lower crust  
473 during continental rifting and subsequent opening of the Rheic ocean. If magmatic rocks of  
474 Cambro-Ordovician age are relatively common in the western part of the Variscan belt,  
475 evidence for metamorphic events of the same age is poorly documented. Abati *et al.* (1999)  
476 reported zircon and monazite U/Pb ages between 484 and 500 Ma for the bimodal magmatism  
477 and the HT metamorphism recorded in the uppermost unit of the Ordenes Complex (NW  
478 Iberia). Monazite U/Pb ages between  $498 \pm 2$  and  $484 \pm 2$  Ma for the metapelitic gneisses have  
479 been interpreted to record an upper amphibolite- and granulite-facies metamorphic event  
480 shortly after the emplacement of the bimodal magmatic rocks (Abati *et al.*, 1999). Moreover, a

481 metapelitic granulite-facies xenolith hosted by the Monte Castello Gabbro gave, for the same  
482 sample, monazite U/Pb ages of  $498 \pm 2$  Ma and rutile U/Pb ages of  $391 \pm 2$  and  $382 \pm 2$  Ma,  
483 very similar to our results. More recently, in the Ossa-Morena Complex (SW Iberia), Solis-  
484 Alulima *et al.* (2020) dated granite emplacement at  $481.2 \pm 2.3$  Ma and LP partial melting  
485 associated with ductile deformation at  $477.8 \pm 2.4$  Ma (SHRIMP zircon U/Pb ages). The authors  
486 suggest a common origin for the migmatitic foliation and the La Cardenchoza Pluton  
487 emplacement in an extensional environment related to continental rifting.

488 Eclogites belonging to the same metamorphic unit as the gneisses studied here have yielded  
489 zircon with magmatic characteristics and a lower intercept age of  $487 \pm 12$  Ma (Paquette *et al.*,  
490 2017), interpreted as the age of the gabbroic protolith emplaced in an oceanic environment.  
491 This age is similar to the youngest age recorded in monazite I from the gneisses. This similarity  
492 suggests that the formation of the oceanic gabbros that gave rise to the eclogites occurred at  
493 about the same time as the HT-LP metamorphism and granite intrusion in the continental unit,  
494 although these two events took place in different geodynamic environments. Godard (1983,  
495 2001) first proposed that the eclogites and altered garnet peridotites (birbirites) represent  
496 fragments of a subducted oceanic lithosphere that were later tectonically incorporated into the  
497 continental crust during the Variscan collision. The discovery of HP parageneses in the  
498 surrounding gneisses (Godard, 2001, 2009) indicates that the tectonic mixing between oceanic  
499 and continental crust occurred before or at the same time as their burial in a subduction zone.  
500 Unfortunately, we have no information about the time and context that brought these units  
501 together, before or during subduction. Similar ages for the oceanic protolith of the eclogites and  
502 the HT-LP parageneses of the continental rocks provide a more accurate picture of the  
503 geodynamic environment of the oceanic and continental units at that time.

### 504 *5.3 Eclogite-facies metamorphic event at 390 Ma*

505 The second metamorphic cycle recorded by the studied gneisses occurred under eclogite-  
506 facies conditions ( $T = 700^{\circ}\text{C}$  and  $P > 1.6$  GPa) and gave rise to many pseudomorphic and  
507 coronitic reactions, as HP parageneses developed at the expense of earlier HT parageneses  
508 (Godard, 2009; Fig. 11). So far, the age of the eclogite-facies metamorphism has not been  
509 determined (see section 2.1). The paragneiss and orthogneiss samples record the transition from  
510 HT to HP parageneses via the reaction of monazite I with the HT plagioclase to form monazite  
511 II during eclogite-facies metamorphism. Because of their small size, monazite II crystals are  
512 not easy to analyse, resulting in a U/Pb age with a large uncertainty. However, we consider the  
513 age of  $395 \pm 9$  Ma discussed above as sufficiently robust to represent the age of the HP  
514 metamorphism (Fig. 11).

515 By considering the correlations between the different units exposed in the Ibero-Armorican  
516 arc and their structural juxtaposition, Ballèvre *et al.* (2009, 2014) define different groups of  
517 tectono-metamorphic units which helps to explain the spread of geochronological data obtained  
518 for these units. Like the Champtoceaux unit, the Keramoine and Kergroaz units at Audierne in  
519 the South Armorican domain, and the Cabo Ortegal and Sobrado Mellide units in Iberia, the  
520 Les Essarts unit belongs to the Upper Allochthon with reference to the Galicia section.  
521 According to this model, below the Upper Allochthon would occur the oceanic complexes of  
522 the Middle Allochthon (among which Groix, Dumet and Bois-de-Céné blueschists in the  
523 Armorican Massif; *e.g.*, Godard *et al.*, 2024), and then the Lower Allochthon which includes  
524 eclogite-bearing units, as the Malpica Tui unit in the Iberia Massif and the Cellier unit in the  
525 Armorican Massif. Eclogites from the Lower Allochthon are characterised by a HP-LT  
526 metamorphism during the Devonian (370–355 Ma), whereas metabasites from the Upper  
527 Allochthon are HP-HT eclogites (or granulites), yielding rather older metamorphic ages of  
528 around 400–390 Ma (Ballèvre *et al.*, 2009, 2014). Moreover, geochemical data in the HT  
529 eclogitic metabasite rocks from the Upper Allochthon show compositions similar to N-MORB,

530 suggesting an oceanic or arc origin (Montigny & Allègre, 1974; Godard, 1983, 1988, 2001;  
531 Bernard-Griffiths & Cornichet, 1985; Gil Ibarra *et al.*, 1990; Galàn & Marcos, 1997). By  
532 comparison, LT eclogites from the Lower Allochthon are thought to be derived from former  
533 dyke swarms intruded into a thinned continental crust (Ballèvre *et al.*, 1994), where  
534 orthogneisses with the same reaction microstructures as the Grezay metagranite studied here  
535 were metamorphosed under eclogite-facies conditions (*e.g.*, Gyomlai *et al.*, 2023). In the Upper  
536 Allochthon, the HP-HT eclogites of the Cabo-Ortegal complex have been dated by an in-situ  
537 method to about 390 Ma (scattered SHRIMP U/Pb ages: Ordonez-Casado *et al.*, 2001). In NW  
538 Portugal, the Bragança eclogites have yielded zircon and rutile U/Pb ages of  $390 \pm 4$  Ma (ID-  
539 TIMS, concordant ages; Roger & Matte, 2005). In the Armorican Massif, HP granulites from  
540 the Audierne region (Kergoaz unit) have given zircon ages of  $384 \pm 6$  Ma (ID-TIMS: Paquette  
541 *et al.*, 1985). Finally, migmatitic gneisses from the Champtoceaux Unit, containing eclogite  
542 lenses, have been dated to about 390 Ma (EMP Th-U/Pb monazite dating; Cocherie *et al.*, 2005),  
543 which probably reflects the age of the partial melting, so that the age of the eclogite-facies  
544 metamorphism in this unit must be older than 390 Ma (Ballèvre *et al.*, 2009). In the French  
545 Massif Central, Sm-Nd dating of the La Bessenoits eclogite has yielded an isochron age of  $408$   
546  $\pm 7$  Ma (Paquette *et al.*, 1995). Recent geochronological studies have given ages in the same  
547 range: garnet Lu/Hf and Sm/Nd ages of  $383 \pm 1$  Ma and  $377 \pm 3$  Ma, respectively, in the Najac  
548 eclogite (Lotout *et al.*, 2018); in-situ U/Pb zircon rim of  $379 \pm 6$  Ma and younger ages with  
549 garnet Lu/Hf dating at  $357 \pm 13$  Ma and  $358 \pm 2$  Ma in the Levezou eclogites (Lotout *et al.*,  
550 2020); in-situ U/Pb zircon rim of around 380–364 Ma (Berger *et al.*, 2010; Benmammar, 2021)  
551 in some Limousin eclogites.

552 These studies show that the HP metamorphic event in the Upper Allochthon is most likely  
553 Lower Devonian in the Iberian-Armorican Arc, and the age of  $395 \pm 9$  Ma obtained for monazite  
554 II from the gneisses of Les Essarts Unit is consistent with this conclusion. As pointed out by

555 Ballèvre *et al.* (2014), the HP-HT eclogite metamorphism at 400–390 Ma together with the  
556 generation of intraoceanic supra-subduction zone ophiolites at 405–395 Ma in the Upper  
557 Allochthon, reflects the first evidence of convergence in the Variscan cycle.

#### 558 *5.4 Consequences for the understanding of the monazite and xenotime chronometers*

559 Monazite is known to be a good tracer of petrological events and a robust geochronometer  
560 because of its high Th and, to a lesser extent, U contents, which allow the simultaneous use of  
561 three isotopic systems as independent radiometric clocks that control each other (Williams *et*  
562 *al.*, 2007; Bosse & Villa, 2019 and references therein). The low diffusion rate of Pb in the  
563 monazite lattice (*e.g.*, Cherniak *et al.*, 2004; Gardès *et al.*, 2006) prevents resetting, even at high  
564 temperatures, and its high capability to recrystallise during interactions with fluids and/or  
565 during deformation (Wawrzenitz *et al.*, 2012; Didier *et al.*, 2014) makes it a useful  
566 geochronometer in polycyclic metamorphic rocks. The present study illustrates these  
567 peculiarities very well: on the one hand, the crystallisation of monazite II during the HP  
568 metamorphic event, at the expense of HT plagioclase and monazite I, provides direct evidence  
569 for the metamorphic reactions taking place in the gneisses during a new orogenic cycle. On the  
570 other hand, the limited volume diffusion within the monazite crystal lattice during the *P-T*  
571 evolution of the rocks allows the preservation of the ages of the early HT event as well as the  
572 record of the HP event. Monazite I isolated from plagioclase shows no textural evidence of  
573 reaction and therefore does not record the Early Devonian HP event. The same is observed for  
574 xenotime, which is not involved in any reaction. Only the rims of the monazite grains at contact  
575 with plagioclase in some samples (VP7-2 and VP7r) are reactive and record the latest  
576 metamorphic event. The small-scale processes occurring at the rims are difficult to estimate.  
577 From Figure 9c-d, we can see that even in the reaction rims apparent (concordant) dates similar  
578 to those obtained from the core of the grains have been measured. Are these relict domains of  
579 monazite I or disturbed apparent dates without geological significance? We cannot answer this



580 question in the present case. Unfortunately, current analytical performances do not yet provide  
581 sufficient spatial resolution to date small grains with a high degree of accuracy. However, this  
582 study shows that the combination of a detailed petrological study and in-situ dating techniques  
583 using robust geochronometers can allow complex and superimposed geological events to be  
584 approached with sufficient precision.

## 585 **6- Conclusions**

586 The Les Essarts unit (Armorican Massif, Vendée, France) contains eclogite km-sized lenses  
587 within orthogneiss and paragneiss that have preserved traces of an early HT stage (granite  
588 intrusion and migmatization). The monazite and xenotime present in the gneisses make it  
589 possible to date the various stages of the polycyclic metamorphic evolution of the unit.

590 In the three samples studied, monazite I occurs either as inclusions in garnet I (related to the  
591 HT event), quartz or biotite in the form of pristine crystals, or at contacts with plagioclase  
592 pseudomorphs. In the latter case, monazite I shows complex coronas composed of apatite,  
593 zoisite and clusters of small grains of monazite II. Like most minerals of the HT or initial  
594 granitic paragenesis, monazite in contact with plagioclase participates in the transition from the  
595 HT to the HP event, and therefore records the age of the HP eclogite-facies metamorphism.

596 In the VP7c paragneiss sample, monazite I gives a lower intercept age of  $479 \pm 4$  Ma, but  
597 the concordant U/Pb dates are scattered between 505 and 462 Ma. No clear coronas with  
598 monazite II were observed in this sample. Monazite I from paragneiss VP7-2 also shows  
599 scattered dates and gives a weighted mean age of  $^{206}\text{Pb}/^{238}\text{U}$  of  $492 \pm 4$  Ma. Considering the  
600 statistical distribution of Th/Pb dates, two peaks appear at similar ages, around 495 and 480  
601 Ma, in the two paragneiss samples, in contrast to the unimodal age distribution for monazite I  
602 from the VP7r orthogneiss. This age spread is interpreted as monazite crystallisation and/or re-  
603 equilibration over a longer period during the prograde phase of the HT cycle. Xenotime is

604 almost exclusively present as inclusions in garnet I of the VP7-2 paragneiss and records the  
605 prograde stage of the HT cycle.

606 Monazite II consists of small crystals that could only be analysed in paragneiss sample VP7-  
607 2, giving concordant ages of about 395 Ma, which we interpret as the age of the eclogite-facies  
608 event in the Les Essarts unit. This age is consistent with the results obtained in the HP  
609 metamorphic rocks of the Upper Allochthon unit of the Iberian-Armorican Arc (Cabo-Ortega,  
610 Bragança, Audierne and Massif Central).

611 In the orthogneiss sample VP7r, xenotime and monazite I belong to the magmatic granite  
612 paragenesis and indicate the age of the granitic protolith at  $496 \pm 13$  Ma and  $496 \pm 10$  Ma,  
613 respectively. Monazite II crystals were too small to be analysed in this sample.

614 Following Ballèvre *et al.* (2012), HT metamorphism and partial melting are thought to have  
615 occurred in the lower crust during continental rifting and the subsequent opening of the Rheic  
616 Ocean. The formation of the oceanic gabbros that gave rise to the neighbouring eclogites  
617 occurred at about the same time as HT-LP metamorphism and granitic intrusion into the  
618 continental unit.

619 The results presented in this study clearly illustrate the qualities of monazite as a useful  
620 geochronometer and a good petrological tracer.

621

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## 789 **Figure captions**

790 **Figure 1:** Geological map of the Les Essarts metamorphic unit (Armorican Massif, France), modified  
791 from Godard (2001). The inset block diagram schematically shows the different geological formations  
792 and their relationship. All the studied samples come from the Grezay locality.

793 **Figure 2:** Coronitic gneisses from Grezay: (a) undeformed migmatite, with eclogite-facies garnet-  
794 bearing microcoronas (not visible) (modified after Godard, 2009); (b) orthogneiss, with K-feldspar  
795 *augen* (Kfs) and similar microcoronas; (c) micrograph of a paragneiss, showing Grt II+Rt microcoronas



796 around ilmenite (Ilm), pseudomorphs after cordierite (ps Crd) and pseudomorphs after plagioclase (ps  
797 Pl), the latter consists of polycrystalline albite and minute kyanite rods delimiting a honeycomb-like  
798 structure, the cells of which could represent former plagioclase single crystals before the albite-jadeite  
799 transition (plane-polarised light; after Godard, 2009); (d) micrograph of a paragneiss, showing  
800 pseudomorphs after plagioclase (ps Pl), Grt II+Ph microcoronas at the interface between biotite (Bt) and  
801 the former plagioclase (ps Pl), and Grt II overgrowths at the interface between pre-existing garnet (Grt  
802 I) and the former plagioclase; a monazite (Mnz) grain developed a radiohalo in biotite (plane-polarised  
803 light).

804 **Figure 3:** BSE images of monazite grains from the VP7-2 paragneiss. Circles indicate the location of  
805 the LA-ICPMS pits (7 µm diameter) and their corresponding  $^{206}\text{Pb}/^{238}\text{U}$  dates ( $2\sigma$  level). Only consistent  
806 dates in both the  $^{208}\text{Pb}/^{232}\text{Th}$ – $^{206}\text{Pb}/^{238}\text{U}$  and  $^{206}\text{Pb}/^{238}\text{U}$ – $^{207}\text{Pb}/^{235}\text{U}$  systems are shown. (a) and (b)  
807 Monazite I grain in pseudomorph after plagioclase (ps Pl) with its reaction corona composed of apatite  
808 (Ap), zoisite (Zo) and monazite II (Mnz II); (c) and (d) similar coronas in which monazite I has been  
809 completely consumed; (e) pristine monazite I included in quartz (Qz); (f) pristine monazite (Mnz I) in  
810 pseudomorph after cordierite (ps Crd), consisting of Grt II, cryptocrystalline kyanite and quartz; (g)  
811 pristine monazite (Mnz I) included in Grt I. Kln = kaolinite.

812 **Figure 4:** BSE images of monazite grains from the VP7c paragneiss. Circles indicate the location of the  
813 LA-ICPMS pits (7 µm diameter) and their corresponding  $^{206}\text{Pb}/^{238}\text{U}$  dates ( $2\sigma$  level). Only consistent  
814 dates in both the  $^{208}\text{Pb}/^{232}\text{Th}$ – $^{206}\text{Pb}/^{238}\text{U}$  and  $^{206}\text{Pb}/^{238}\text{U}$ – $^{207}\text{Pb}/^{235}\text{U}$  systems are shown. (a) Pristine  
815 monazite I included in garnet I; (b) open inclusion of monazite I in garnet I in contact with biotite (Bt);  
816 note the presence of small inclusions of xenotime (Xtm) in garnet I (see text); (c) pristine monazite I in  
817 a matrix of biotite (Bt)/phengite (Ph); (d) monazite I in plagioclase pseudomorph (ps Pl) showing  
818 irregular corroded rims; the reaction corona at the monazite-plagioclase interface is not observed in this  
819 sample; (e) and (f) pristine monazite I in a matrix of biotite (Bt)/phengite (Ph).

820 **Figure 5:** BSE images of monazite grains from the VP7r orthogneiss. Circles indicate the location of  
821 the LA-ICPMS pits (7 µm diameter) and their corresponding  $^{206}\text{Pb}/^{238}\text{U}$  dates ( $2\sigma$  level). Only consistent  
822 dates in both the  $^{208}\text{Pb}/^{232}\text{Th}$ – $^{206}\text{Pb}/^{238}\text{U}$  and  $^{206}\text{Pb}/^{238}\text{U}$ – $^{207}\text{Pb}/^{235}\text{U}$  systems are shown; dashed circles  
823 indicate the location of the other spot analyses. (a), (b) and (c) Monazite I grains in plagioclase

824 pseudomorphs (ps Pl) with reaction microcoronas composed of apatite (Ap), zoisite (Zo) and monazite  
825 II (Mnz II; the small crystals could not be analysed).

826 **Figure 6:** Chemical composition of monazite from the three studied samples: (a) Ternary diagram  
827 La+Ce+Nd+Pr – U+Th+Pb – Sm+Gd+Y (mol%) showing the distribution of the monazite composition  
828 amongst the main end-members: monazite *s.s.*, actinide-bearing end-members and mid-heavy REE  
829 phosphates; (b) Y-Th-U ternary diagram showing the monazite chemical composition in relation to the  
830 textural position in each studied sample; (c) chondrite-normalised REE patterns in monazite from the  
831 paragneiss sample VP7-2.

832 **Figure 7:** Images of xenotime from the VP7-2 paragneiss (a-d) and VP7r orthogneiss (e, f). BSE images  
833 (b-f) and RGB image (a) with channel intensities correlated to Fe (red), Ca (green) and Mn (blue)  
834 contents. Circles indicate the location of the LA-ICPMS pits (7 µm diameter) and their corresponding  
835  $^{206}\text{Pb}/^{238}\text{U}$  dates (2σ level). (a) and (b) Garnet I with an overgrowth of Ca-rich garnet II (Grt II) at the  
836 contact with the plagioclase pseudomorph (ps Pl); note that xenotime inclusions are only present in  
837 garnet I (see text); (c) and (d) xenotime inclusions in garnet I and in the matrix of the paragneiss sample  
838 VP7-2; note the presence of small xenotime inclusions (Xtm) in the garnet core (Grt I); (e) and (f)  
839 subhedral xenotime grains (Xtm) in the matrix of the orthogneiss sample VP7r; dashed circles indicate  
840 the location of the discordant spot analyses.

841 **Figure 8:** Xenotime geochronological results for the VP7-2 paragneiss. All data are uncorrected for  
842 common Pb. Tera-Wasserburg diagram showing the U/Pb results from xenotime. White ellipses  
843 represent discordant data.

844 **Figure 9:** Monazite geochronological results for the VP7-2 (a-d) and VP7c (e, f) paragneisses. All data  
845 are uncorrected for common Pb. (a) Tera-Wasserburg diagram showing the U/Pb results for monazite I.  
846 White ellipses correspond to discordant analyses. White dashed ellipses are concordant but were not  
847 included in the weighted mean dates calculated from the grey ellipses. (b) Probability histogram of  
848  $^{208}\text{Pb}/^{232}\text{Th}$  dates for monazite I (only concordant data in both the U/Pb systems and the  $^{206}\text{Pb}/^{238}\text{U}$  –  
849  $^{208}\text{Pb}/^{232}\text{Th}$  systems). (c) Tera-Wasserburg diagram showing the U/Pb results for monazite II. Only the  
850 dark grey ellipses were taken into account in the free regression calculation. In the inset (concordia date),  
851 the light grey ellipse with the bold outline corresponds to the calculated concordia date. (d) Probability

852 histogram of  $^{208}\text{Pb}/^{232}\text{Th}$  dates for monazite II (only concordant data in both the U/Pb systems and the  
853  $^{206}\text{Pb}/^{238}\text{U}$ – $^{208}\text{Pb}/^{232}\text{Th}$  systems). (e) Tera-Wasserburg diagram showing the U/Pb results in monazite I.  
854 White ellipses have not been considered in the free regression calculated with the grey ellipses. (f)  
855 Probability histogram of  $^{208}\text{Pb}/^{232}\text{Th}$  dates for monazite I. Only consistent dates in both the  $^{208}\text{Pb}/^{232}\text{Th}$ –  
856  $^{206}\text{Pb}/^{238}\text{U}$  and  $^{206}\text{Pb}/^{238}\text{U}$ – $^{207}\text{Pb}/^{235}\text{U}$  systems are shown in (b), (d) and (f).

857 **Figure 10:** Monazite and xenotime geochronological results for the VP7r orthogneiss. All data are  
858 uncorrected for common Pb. (a) Tera-Wasserburg diagram showing the U/Pb results for monazite I. The  
859 white ellipse has not been considered in the free regression calculated with the grey ellipses. (b) Tera-  
860 Wasserburg diagram showing the U/Pb results for xenotime. The white ellipses have not been considered  
861 in the free regression calculated with the grey ellipses. Only one analysis was consistent in both the  
862  $^{208}\text{Pb}/^{232}\text{Th}$ – $^{206}\text{Pb}/^{238}\text{U}$  and  $^{206}\text{Pb}/^{238}\text{U}$ – $^{207}\text{Pb}/^{235}\text{U}$  systems, so the probability histogram is not available  
863 for xenotime in this sample.

864 **Figure 11:** P-T-time path recorded in the VP7 gneisses of Grezay (Les Essarts unit; based on Figure 16  
865 of Godard, 2009). The age range between 510 and 480 Ma is interpreted as monazite I crystallisation  
866 and/or re-equilibration during the prograde phase and peak of the HT cycle (paragneiss samples VP7c  
867 and VP7-2). Granite emplacement is dated to about 496 Ma in sample VP7r. The HP eclogite-facies  
868 event is dated to about 395 Ma by monazite II from the VP7-2 paragneiss. The two BSE images  
869 correspond to monazite I included in garnet I of the VP7c paragneiss (Fig. 4a) and crystallized during  
870 the HT metamorphic cycle, and to monazite II from the reaction corona formed during the eclogite-  
871 facies HP metamorphism at the monazite I-plagioclase interface in the VP7-2 paragneiss (Mnz VP7-2-  
872 15; SM2).

873

#### 874 **Supplementary materials**

875 SM1: EMP analyses in monazite (wt.% and cationic formula for 4 O)

876 SM2: Th-U/Pb isotopic data of monazite and xenotime

877 SM3: Trace element analyses of monazite from the sample VP7-2

878 SM4: Analytical methods;

879 SM5: X-ray element maps of monazite crystals. The maps were acquired at the EPMA by wavelength

880 dispersive spectrometry of Ce ( $L\alpha$  line), Gd ( $L\beta$  line), Th ( $M\alpha$  line) and Y ( $L\alpha$  line). Monazite crystals  
881 VP7C-40, VP7C-45, VP7-2-12 (all from paragneiss) are pristine monazite (Mnz I) isolated in garnet  
882 (Grt I) or quartz (Qz), showing in some cases Th-rich inherited nuclei (VP7C-45, VP7-2-12). Monazite  
883 crystals VP7-2-3, VP7-2-7 (from paragneiss), VP7r-1 and VP7r-8 (from orthogneiss) reacted with the  
884 hosting plagioclase (ps Pl) to form a Mnz II-bearing corona around a core of preserved Mnz I. Monazite  
885 II commonly shows enrichment in Th and Y relative to Mnz I (VP7r-1 and VP7r-8). See also Figures.  
886 3e (VP7-2-12), 3a (VP7-2-3), 4a (VP7r-1), 4c (VP7r-8).

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